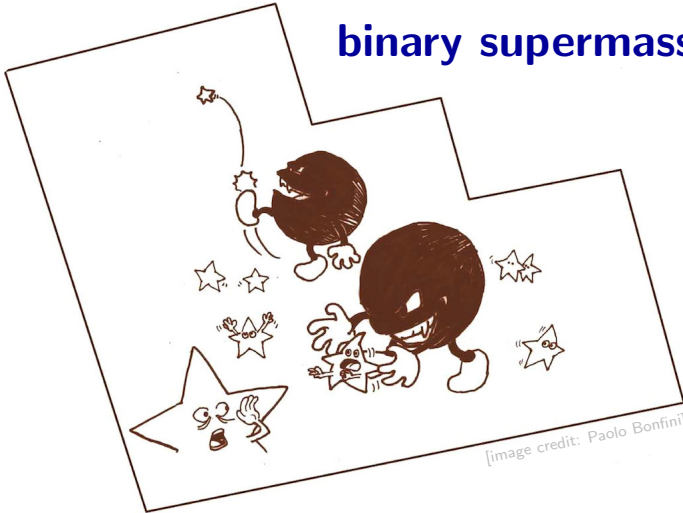


# Life and adventures of binary supermassive black holes

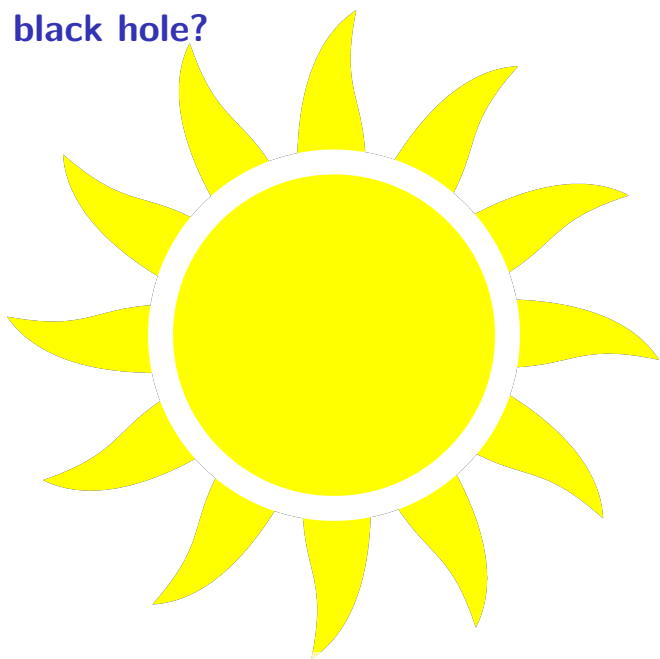


Eugene Vasiliev

Edinburgh, December 2023

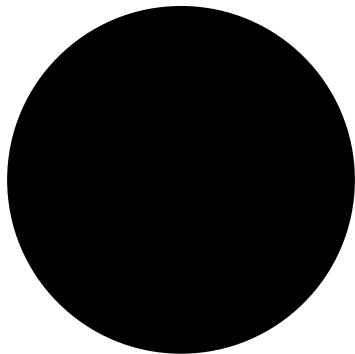
HST captures a binary SMBH kicking stars out of a galaxy

What is... a black hole?



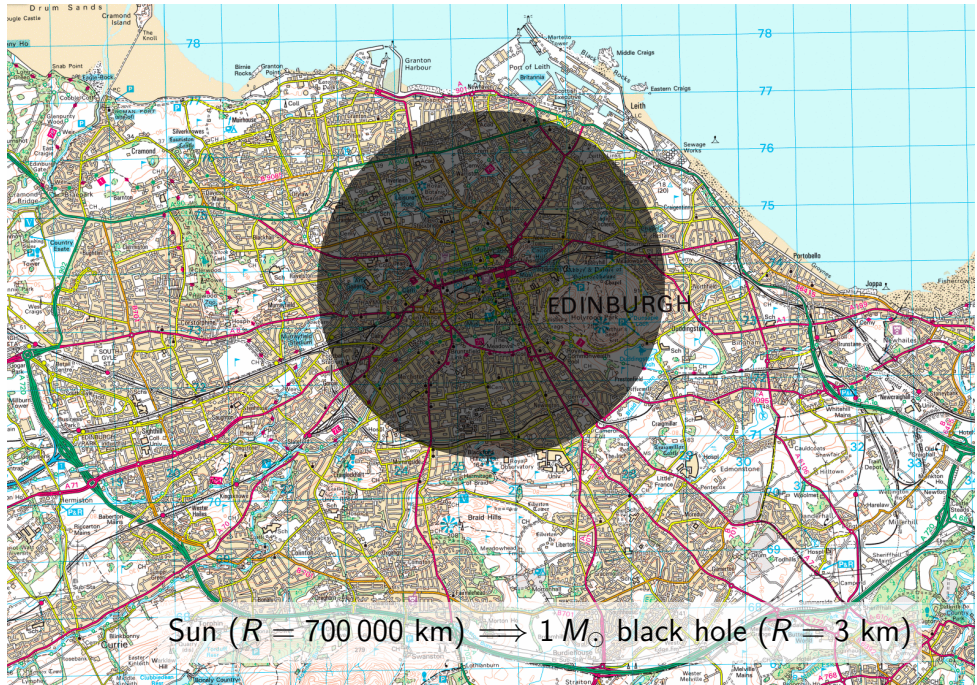
Sun ( $R = 700\,000$  km)

## What is... a black hole?

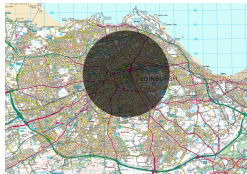


Sun ( $R = 700\,000$  km)  $\implies$   $1 M_{\odot}$  black hole ( $R = 3$  km)

# What is... a black hole?



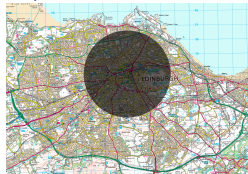
## What is... a supermassive black hole (SMBH)?



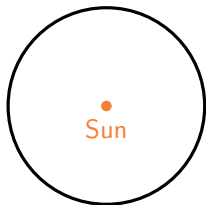
$$1 M_{\odot} \Rightarrow R = 3 \text{ km}$$

for black holes,  $R \propto M$

# What is... a supermassive black hole (SMBH)?



$1 M_{\odot} \Rightarrow R = 3 \text{ km}$



Sagittarius A\*  
(Milky Way centre)

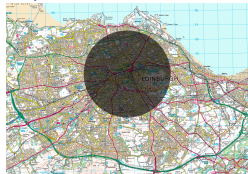
$M = 4.3 \text{ million } M_{\odot}$

$R = 13 \text{ million km}$

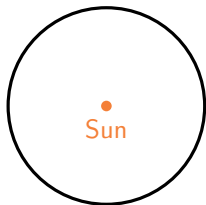
for black holes,  $R \propto M$

Mercury's orbit

# What is... a supermassive black hole (SMBH)?

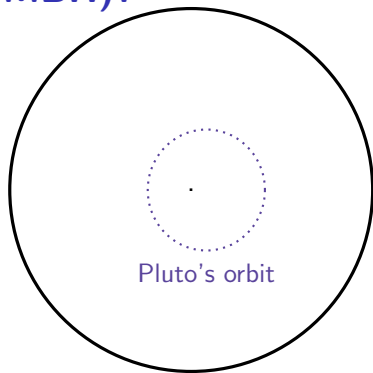


$1 M_{\odot} \Rightarrow R = 3 \text{ km}$



Sagittarius A\*  
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$M = 4.3 \text{ million } M_{\odot}$   
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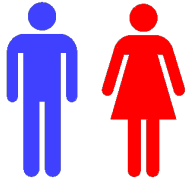
black hole in M87  
(central galaxy of the Virgo cluster)

$M = 6 \text{ billion } M_{\odot}$   
 $R = 18 \text{ billion km}$

for black holes,  $R \propto M$

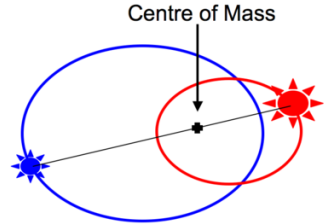
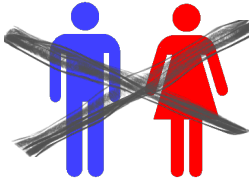
Mercury's orbit

What is... a binary SMBH?





# What is... a binary SMBH?

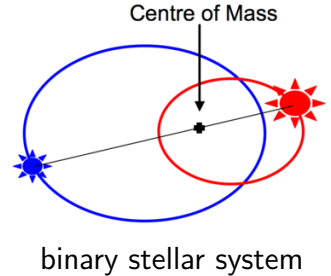
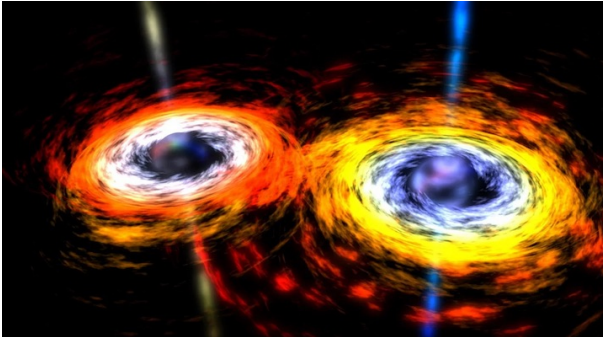


binary stellar system

# What is... a binary SMBH?

– two SMBHs orbiting each other!

(and not just roaming somewhere in the same galaxy)



# What black holes are and are not?



cosmic villains



vacuum cleaners

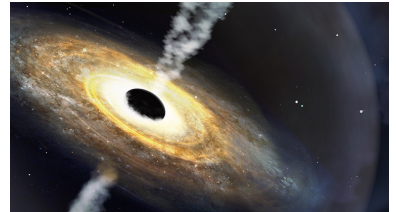
# What black holes are and are not?



cosmic villains



vacuum cleaners

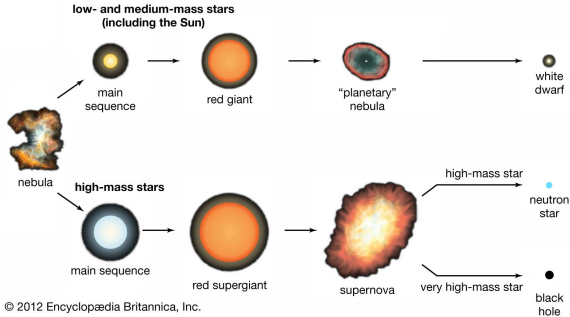


gravitating masses

(possibly accreting surrounding gas,  
which is heated by friction)

# How do black holes form?

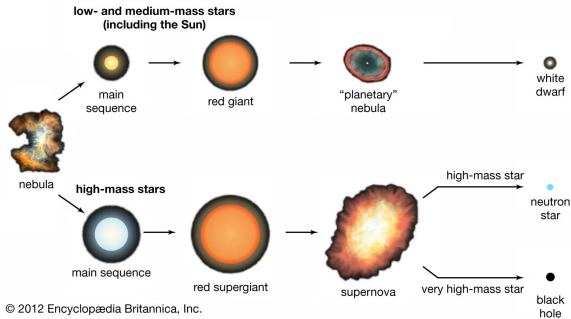
## Stellar evolution



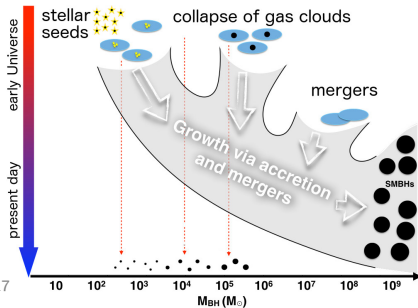
Stellar black holes  
( $M \simeq 5 - 100 M_{\odot}$ )

# How do black holes form?

## Stellar evolution

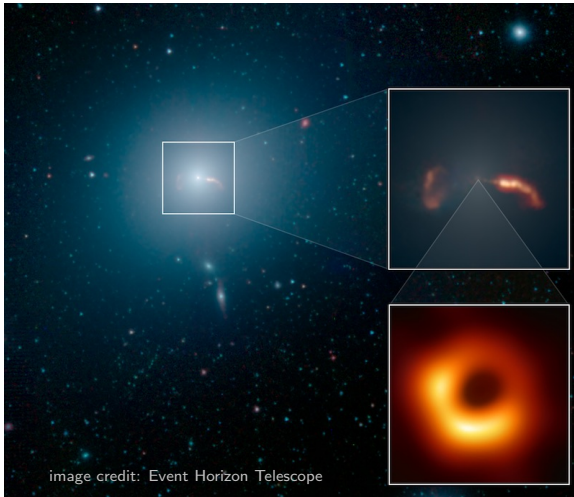


Stellar black holes  
( $M \simeq 5 - 100 M_{\odot}$ )



Supermassive black holes  
( $M \simeq 10^5 - 10^{10} M_{\odot}$ )

## Where do supermassive black holes (SMBH) live?

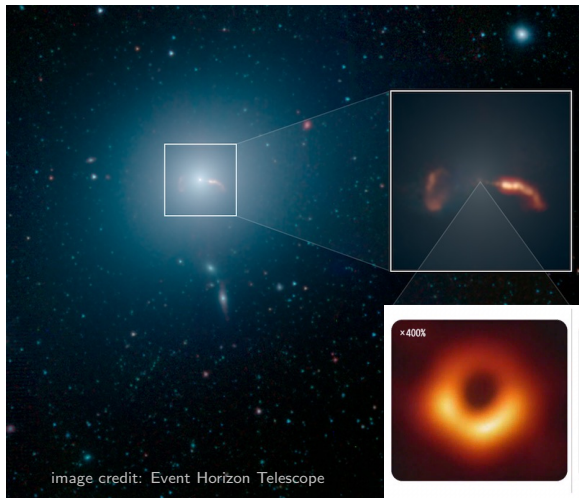


M 87

SMBH mass:  $6 \times 10^9 M_{\odot}$

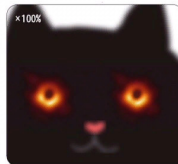
SMBHs are found at the centres of most galaxies!

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M 87

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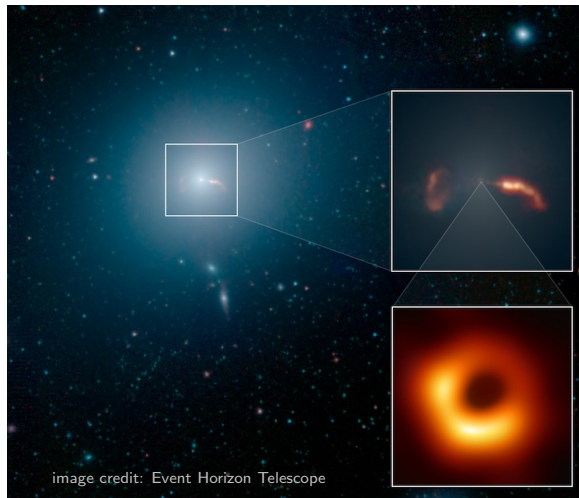


SMBHs are found

ies!

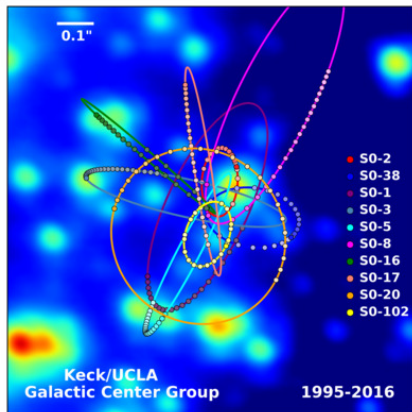


# Where do supermassive black holes (SMBH) live?



M 87

SMBH mass:  $6 \times 10^9 M_{\odot}$

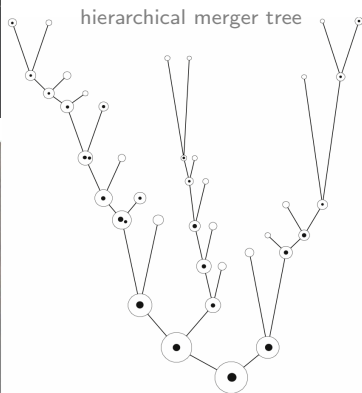
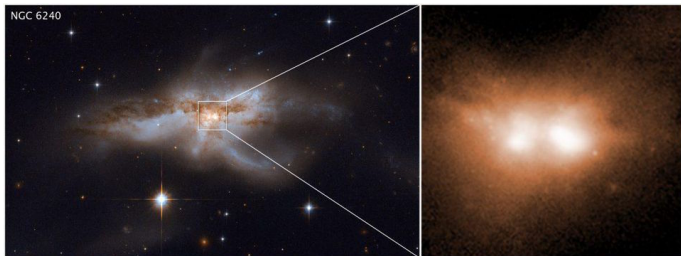
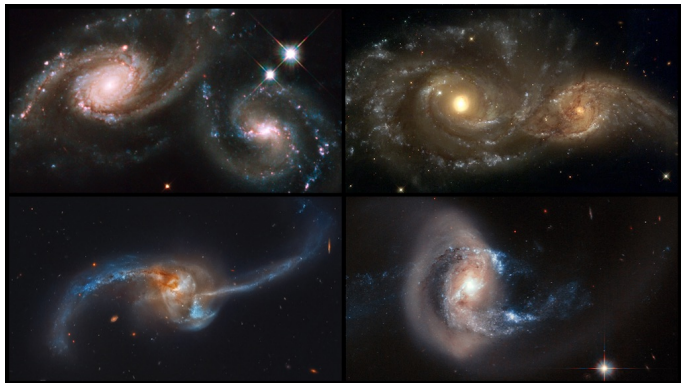


Milky Way (Sgr A<sup>\*</sup>)

mass:  $4 \times 10^6 M_{\odot}$

SMBHs are found at the centres of most galaxies!

# How do binary SMBHs come about?

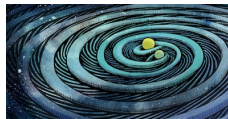
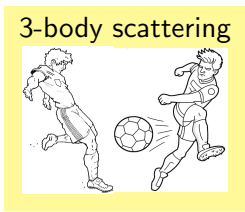
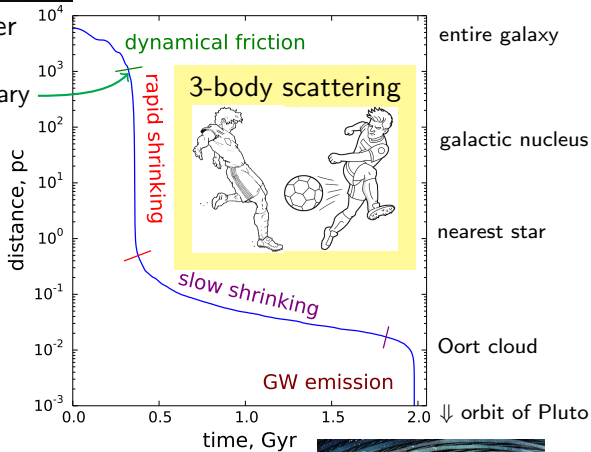


# Life path of a typical binary SMBH



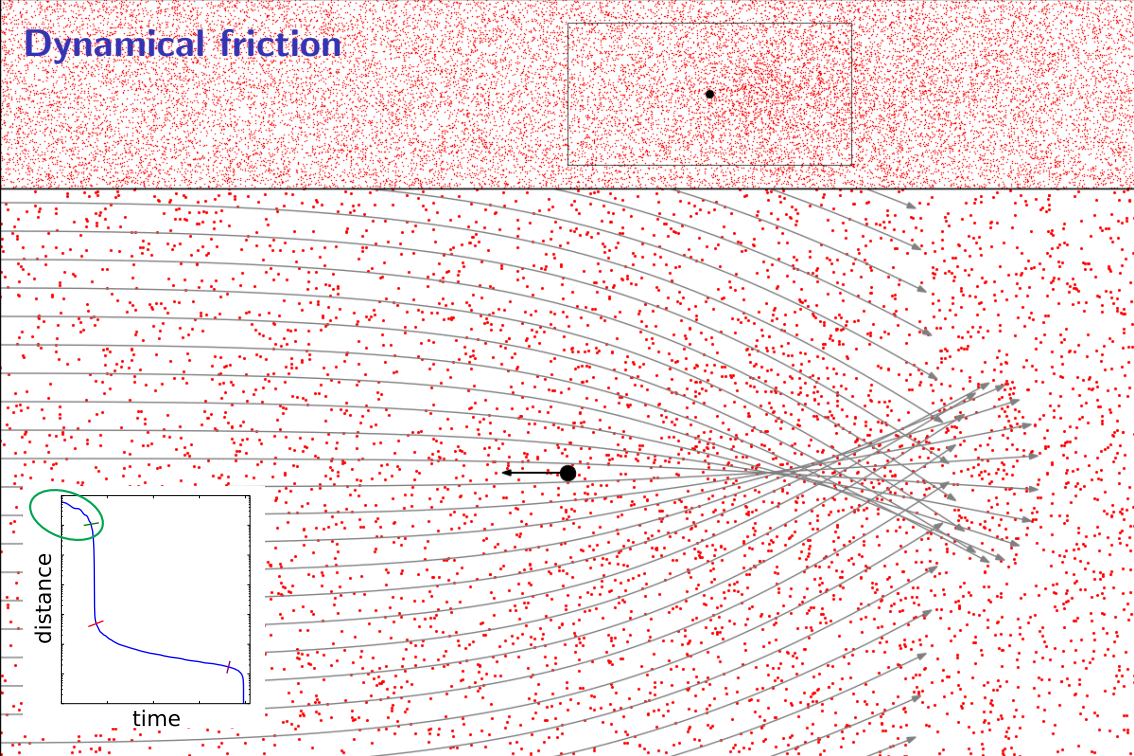
galaxy merger

formation of the binary

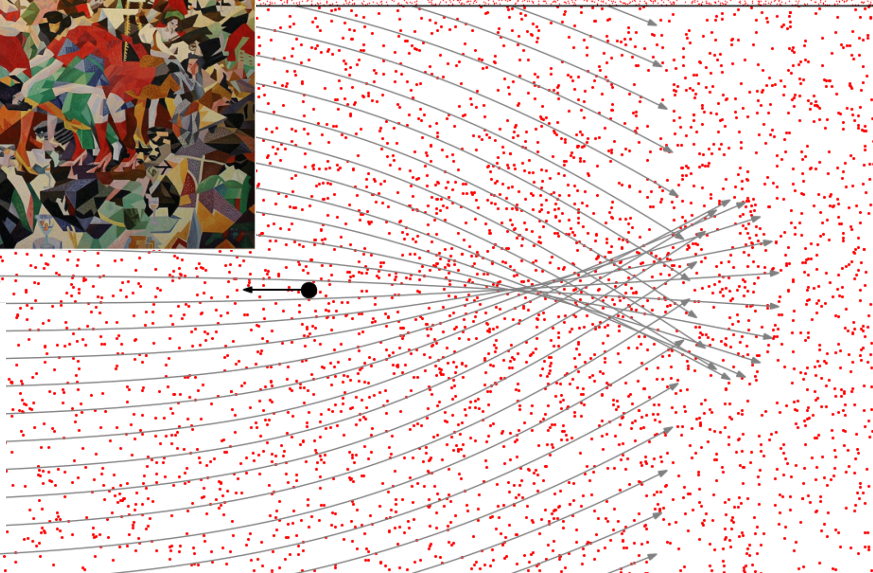
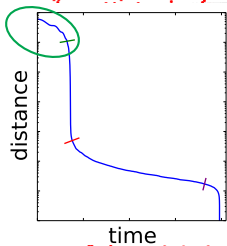
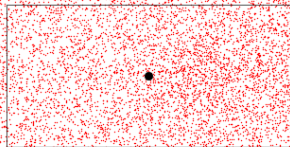


SMBH coalescence

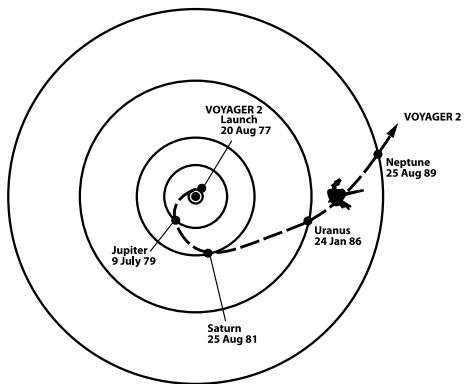
# Dynamical friction



# Dynamical friction



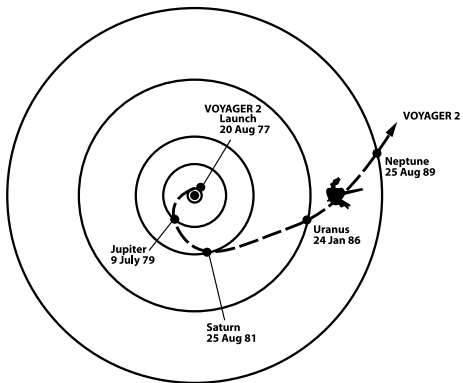
# Gravitational slingshot



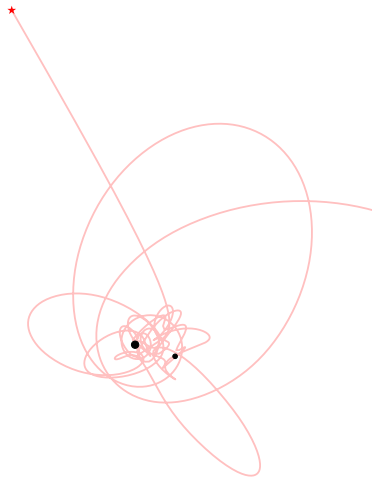
interplanetary transport

# Gravitational slingshot

ejected stars extract energy from the binary SMBH  $\Rightarrow$  its orbit shrinks



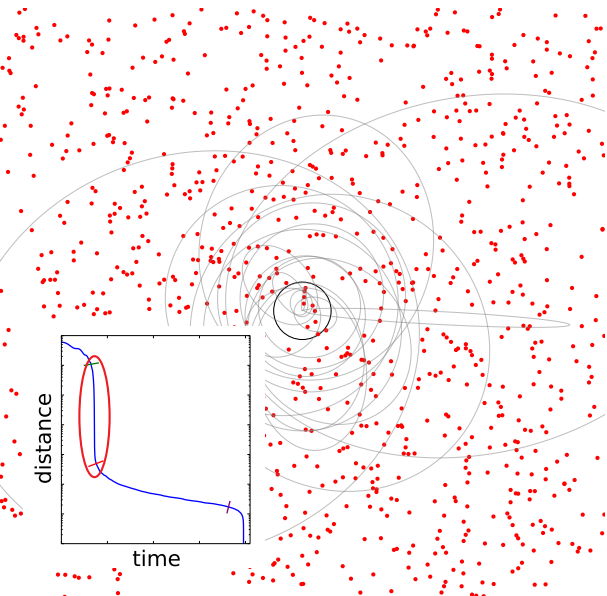
interplanetary transport



3-body scattering

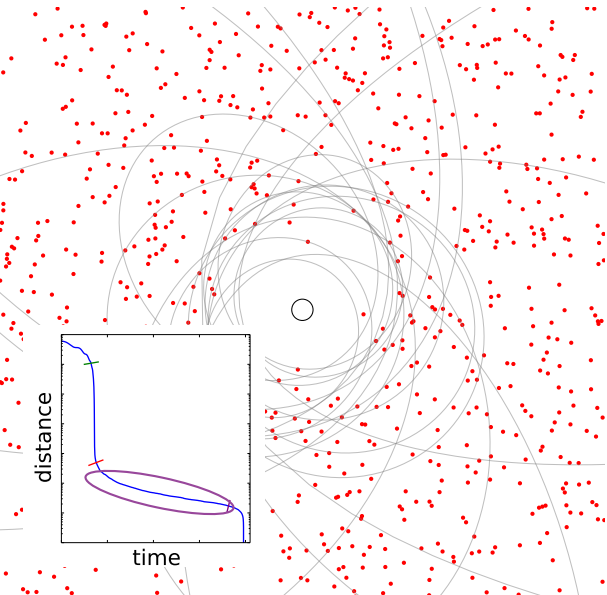


# Shrinking of the binary



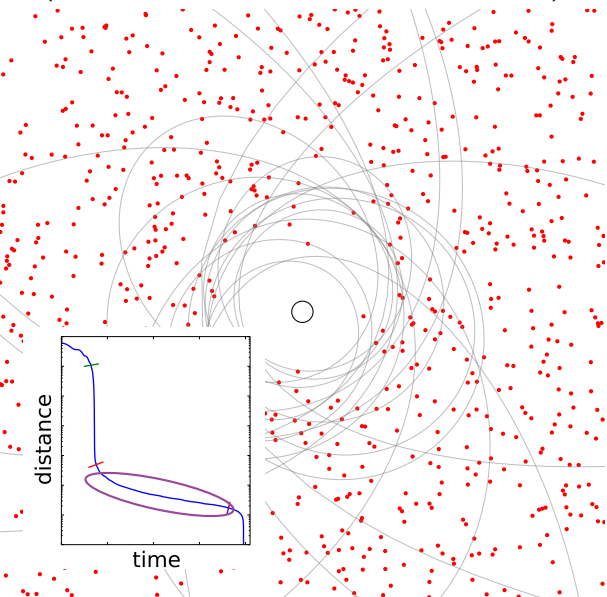


# Shrinking of the binary



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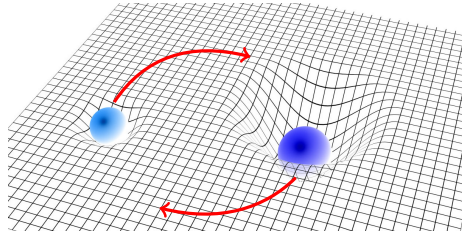
The final-parsec problem!  
(stalling of the binary shrinking at  $\sim 1$  pc)



## Energy loss to gravitational waves

An object moving on a curved trajectory emits gravitational waves, which carry away energy.

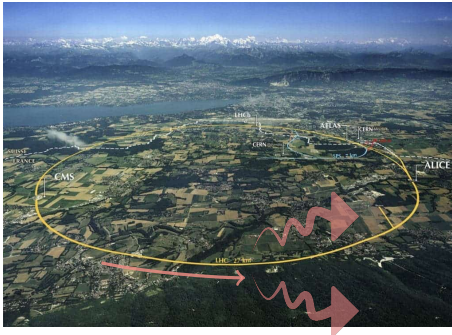
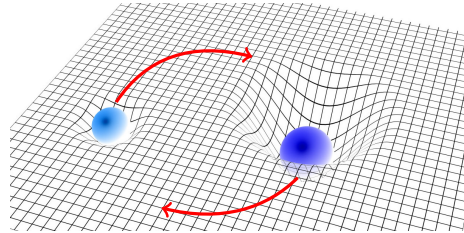
Rate of orbital shrinking rapidly accelerates as the binary orbit becomes smaller and its orbital velocity approaches speed of light.



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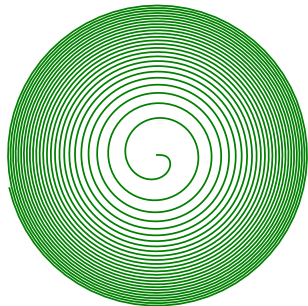
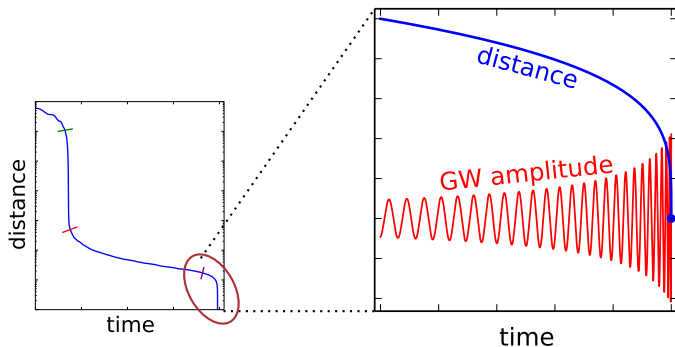
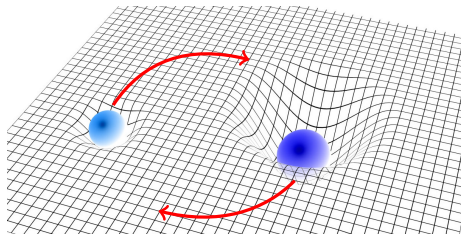


It is analogous to synchrotron radiation emitted by charged particles circulating in the Large Hadron Collider.

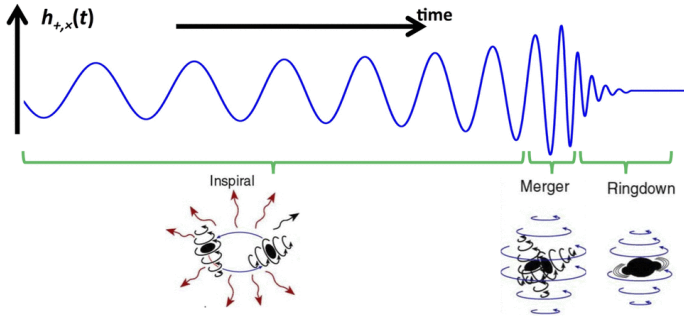
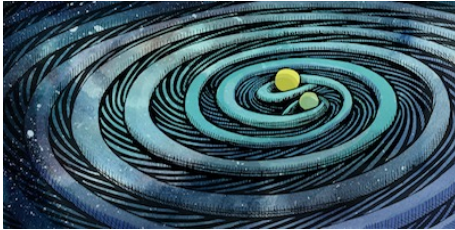
## Energy loss to gravitational waves

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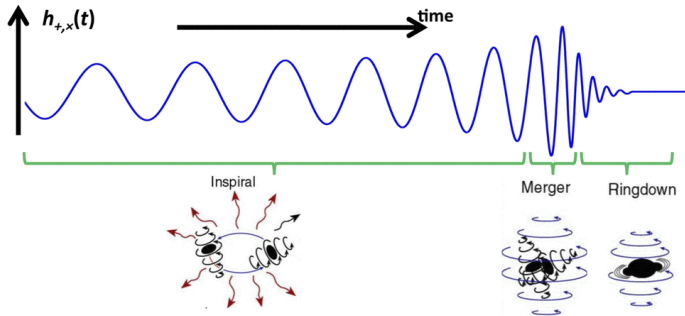
Rate of orbital shrinking rapidly accelerates as the binary orbit becomes smaller and its orbital velocity approaches speed of light.



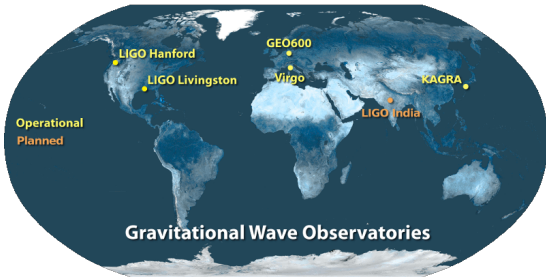
# Final coalescence



# Final coalescence

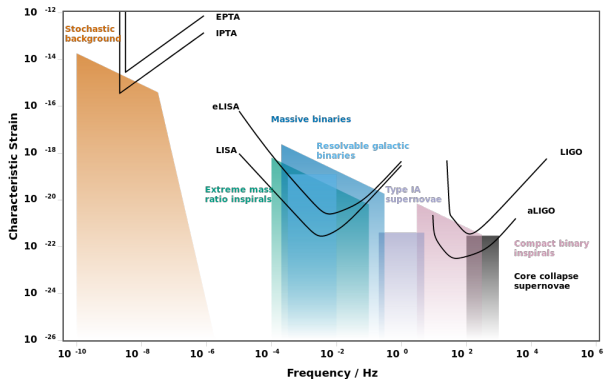


# Gravitational-wave observatories





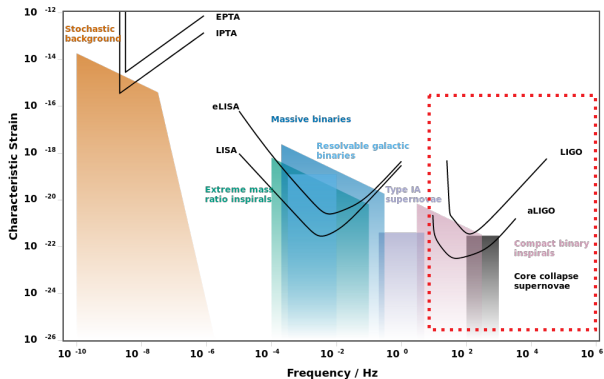
# Gravitational-wave observatories



SMBH  
 $10^6 M_{\odot}$

stellar BH  
 $10 M_{\odot}$

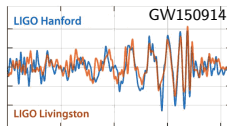
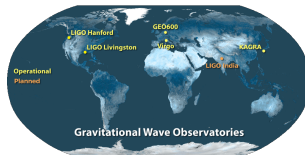
# Gravitational-wave observatories



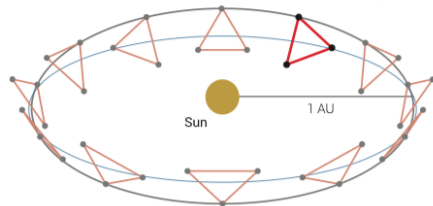
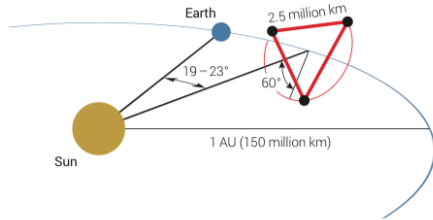
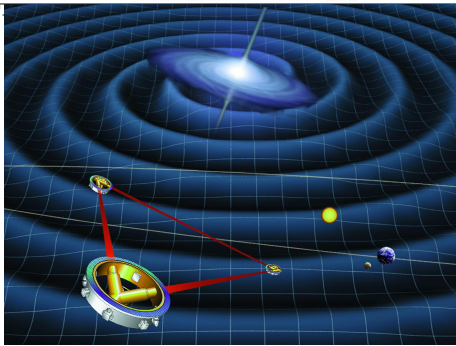
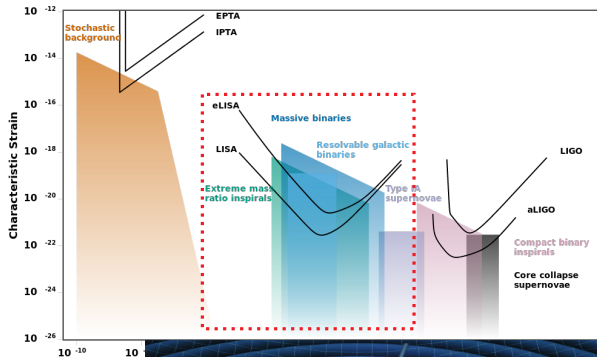
SMBH  
 $10^6 M_{\odot}$

stellar BH  
 $10 M_{\odot}$

## ground-based observatories



# Gravitational-wave observatories in space

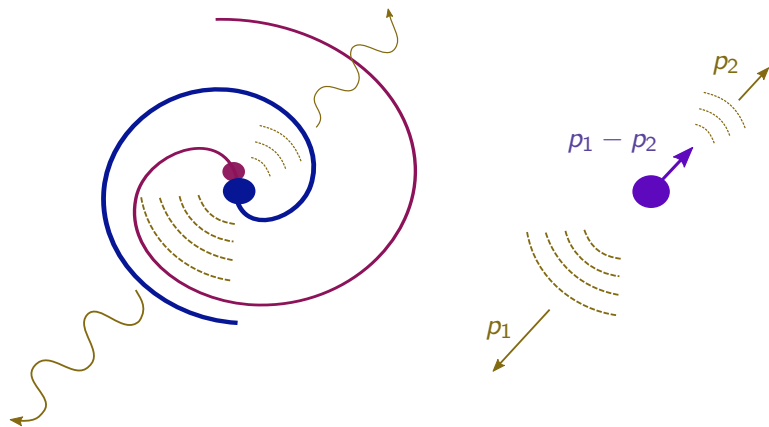


LISA space mission (ESA, 2030s)

## Final handwave

Anisotropic emission of gravitational waves

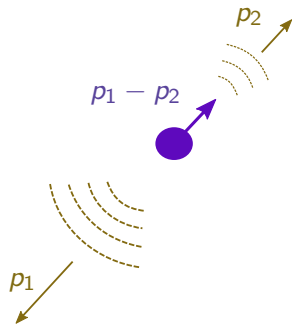
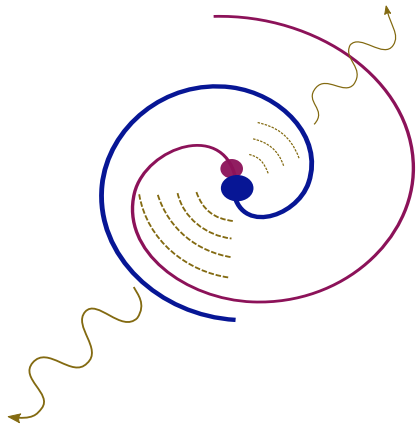
$\implies$  recoil velocity  $\sim 100 - 1000$  km/s



# Final handwave

Anisotropic emission of gravitational waves

$\implies$  recoil velocity  $\sim 100 - 1000$  km/s



# Summary: life cycle of binary SMBHs

